

# Ultrasensitive Infrared Spectroscopy of Molecular Ions

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## Motivations

### Why Molecular Ions?

- Atmospheric Chemistry:** Molecular ions are the key players in the chemical and energetic balance in the ionosphere and auroral regions of Earth's atmosphere and other planetary atmospheres.
- Combustion:** Molecular ions are potential precursors to soot nucleation, and possible initiators of ignition outside of usual combustion conditions.
- Propulsion:** Molecular ions represent attractive fuels for ion engines, as well as promising components of new high energy density materials.
- Astrochemistry:** Molecular ions drive interstellar chemistry.

### Why High-Resolution Spectroscopy?

- Enabling Searches:** Permits searches for molecular ions in the interstellar medium, planetary atmospheres, and combustion environments.
- Benchmarks for Theory:** Spectroscopy is the gold standard for validating quantum chemical calculations.

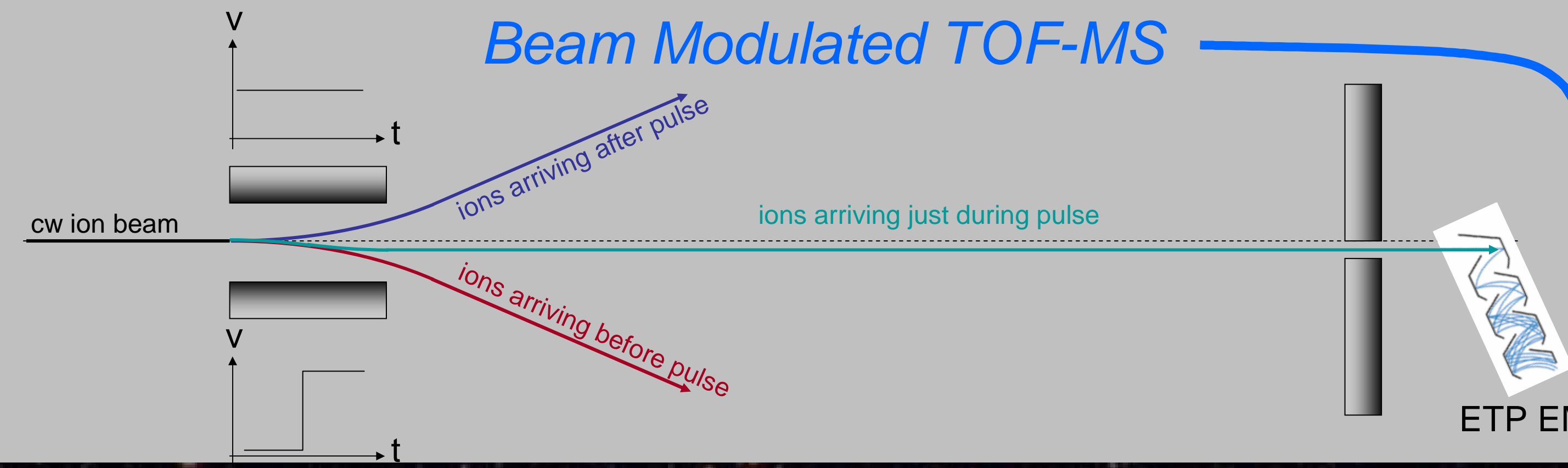
## Ion Optics & Mass Spectrometer

### Ion Optical Design

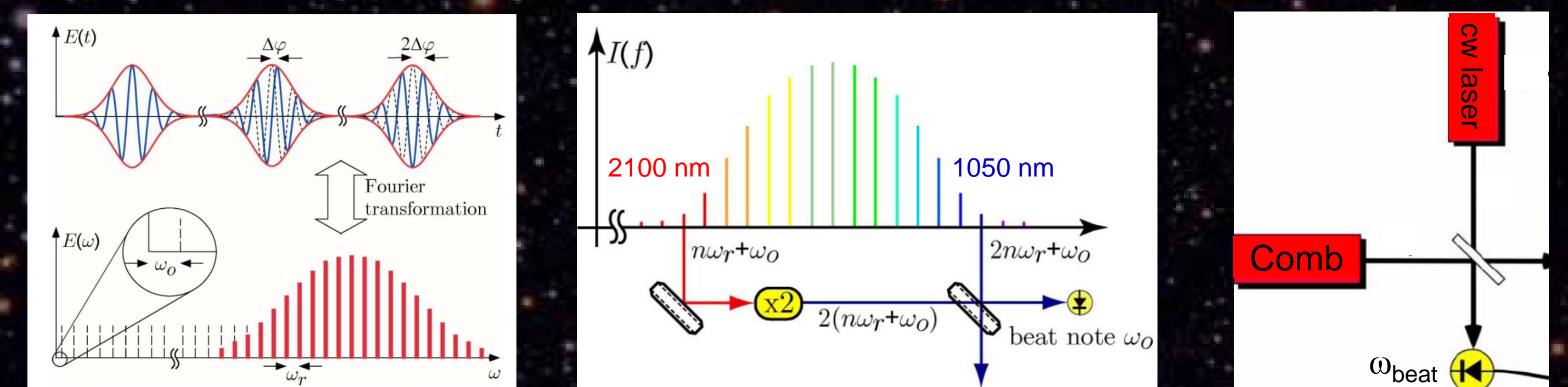


An Einzel lens is used to focus the ion beam from the source... then the beam is steered and "walked" using a set of deflector plates... and a cylindrical bender turns it to overlap with the laser.

### Beam Modulated TOF-MS



## Optical Frequency Comb



A pulse train of a stabilized fs laser is also an optical frequency comb. Doubling the comb, and beating the two combs, yields the offset frequency. Beating the comb with a cw laser yields a beat frequency.

$$\omega_{Ti} = n_{Ti} \cdot \omega_{rep} + 2\omega_{offset} + \omega_{Ti,beat}$$

$$- [\omega_{YAG} = n_{YAG} \cdot \omega_{rep} + \omega_{offset} + \omega_{YAG,beat}]$$

$$= \omega_{DFG} = \underbrace{(n_{Ti} - n_{YAG}) \omega_{rep}}_{\substack{\text{integer} \\ \sim 800,000}} + \underbrace{\omega_{offset}}_{\substack{100 \text{ MHz} \\ \pm 1 \text{ mHz}}} + \underbrace{\omega_{offset}}_{\substack{\sim 20 \text{ MHz} \\ \pm 2 \text{ Hz}}} + \underbrace{(\omega_{Ti,beat} - \omega_{YAG,beat})}_{\substack{\sim 10 \text{ MHz} \\ \pm 50 \text{ kHz}}}$$

MenloSystems FC1500

## Mid-Infrared Laser System

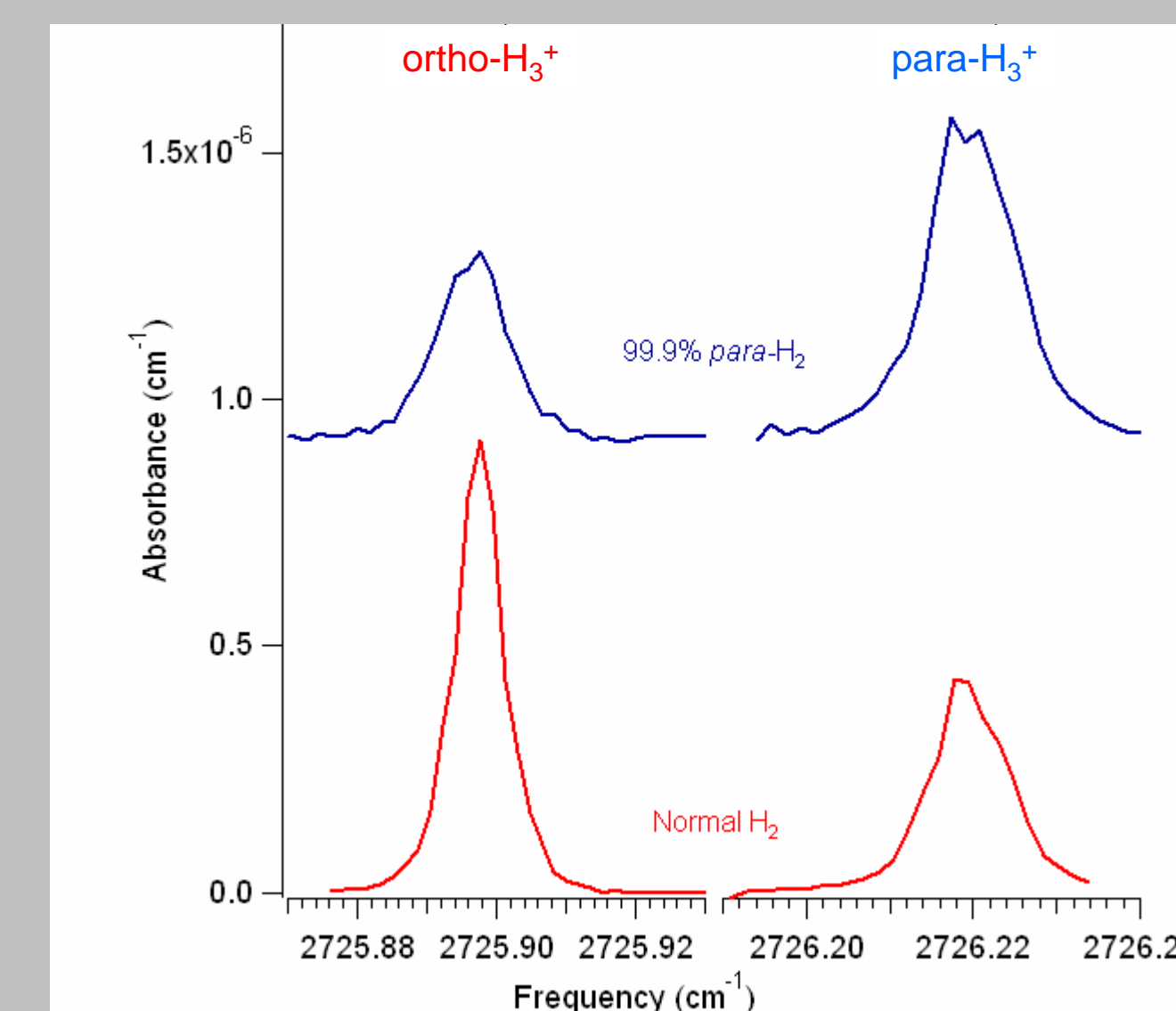
### Difference Frequency Generation



- Tunable 2.8-4.8  $\mu\text{m}$
- Power > 0.5 mW
- Linewidth  $\sim 1$  MHz

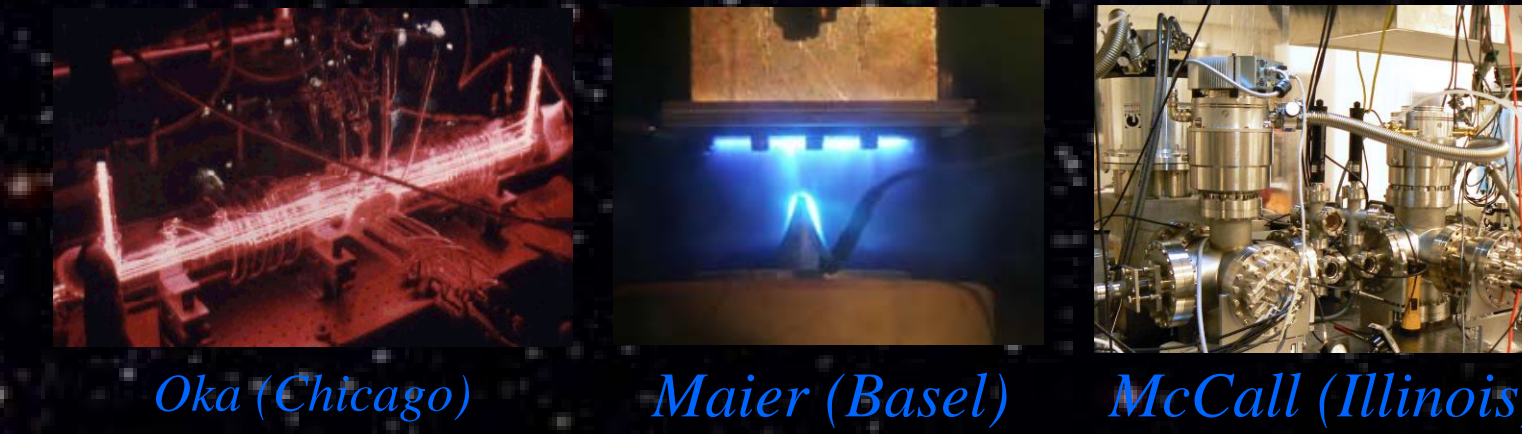
### Initial Spectroscopic Results

Cavity ringdown spectra of  $\text{H}_3^+$  in a supersonic expansion discharge source.



## Comparison with Other Techniques

Two techniques have been widely used for molecular ion spectroscopy, but both have inherent limitations.

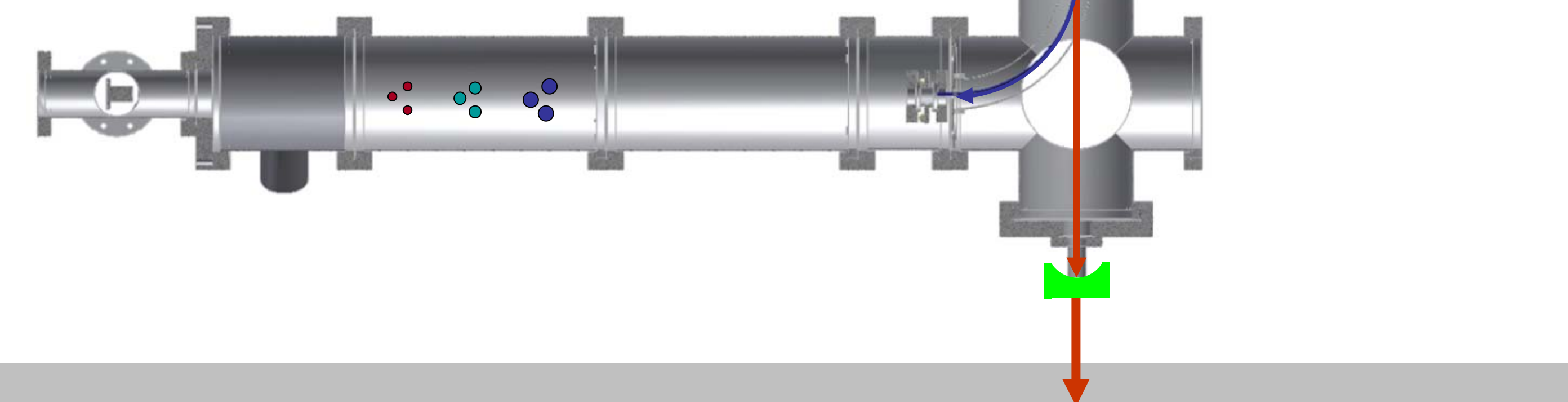


	Velocity Modulation	Supersonic Expansion	SCRIBES
High ion column density	✓		
Ion-neutral discrimination	✓		✓
Low rotational temperature		✓	✓
Sub-Doppler linewidth		✓	✓
Ultra-narrow linewidth (kinematic compression)			✓
Compatible with cavity-enhanced spectroscopy		✓	✓
Mass spectrometry of laser-probed ions			✓
Spectral identification of ion mass			✓

# SCRIBES

Sensitive Resolved Ion Beam Spectroscopy

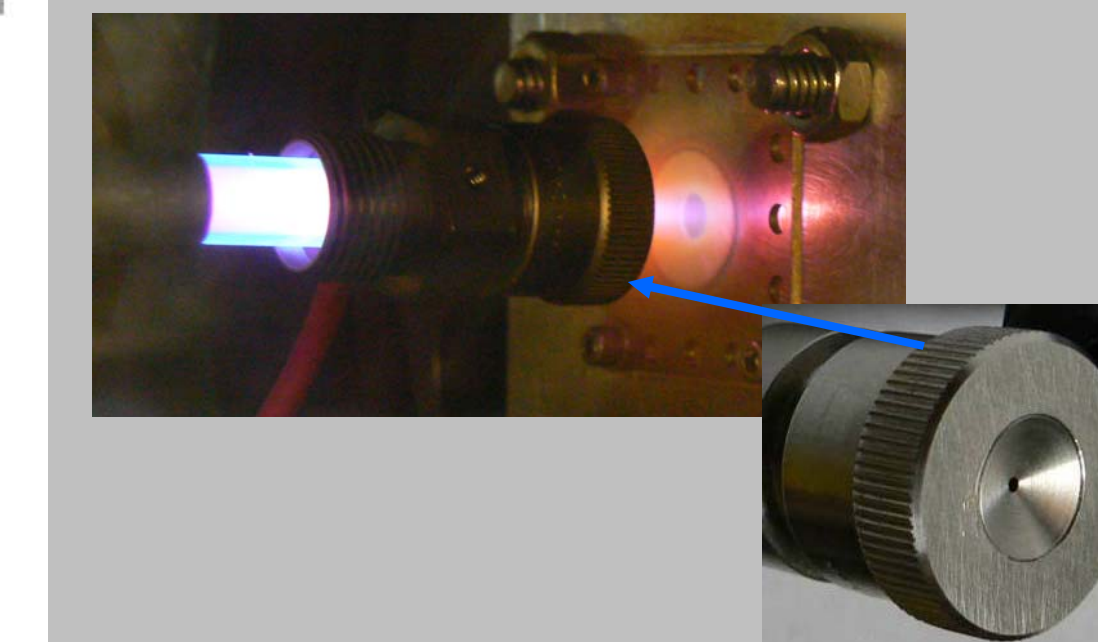
A versatile new technique for high-resolution, low-temperature molecular ion spectroscopy



## Molecular Ion Sources

### Cold Cathode

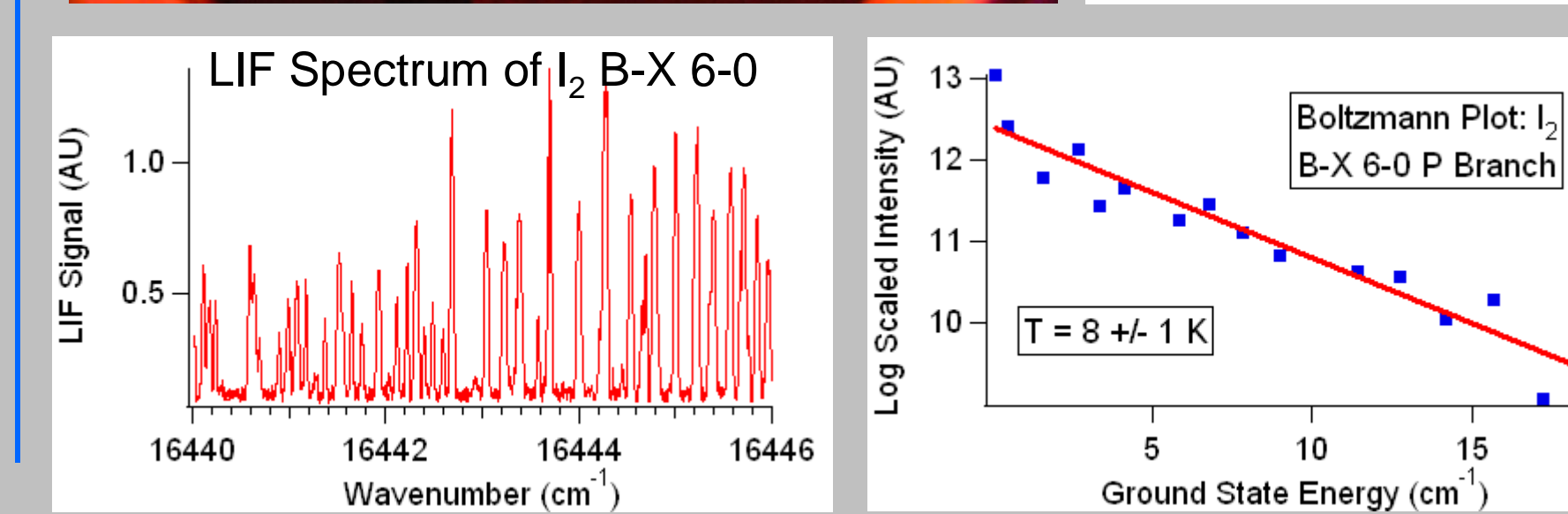
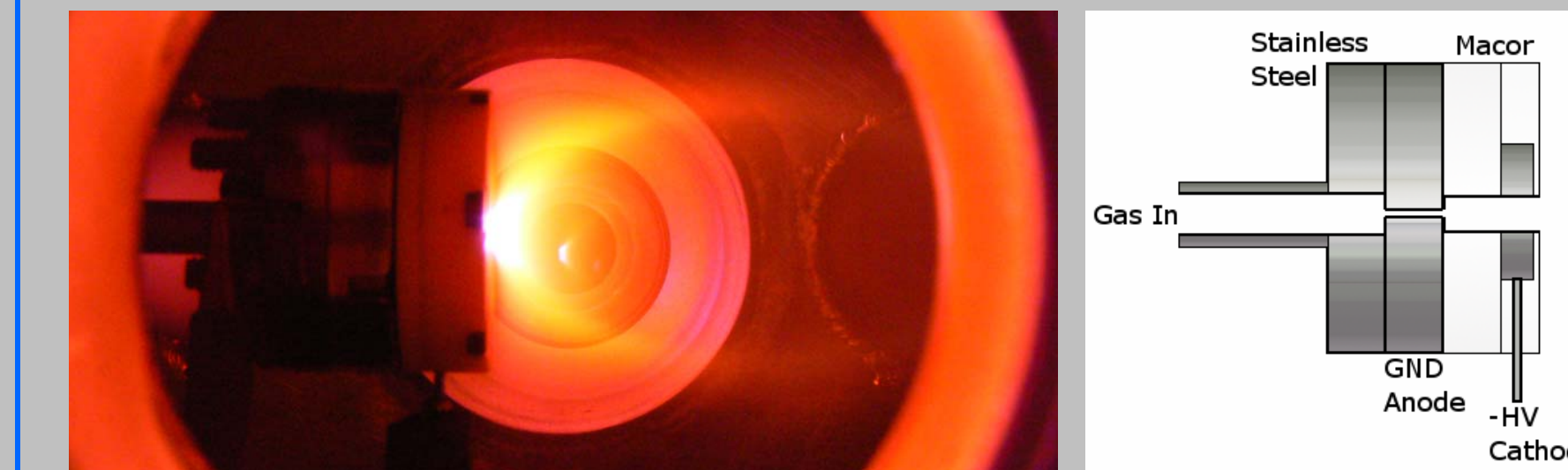
Our initial test source is an uncooled cold cathode (adapted from a Saykally group design).



- Discharge current:  $\sim 1$  mA
- Emission current:  $\sim 50$   $\mu\text{A}$
- Beam current:  $\sim 1.5$   $\mu\text{A}$  ( $\text{N}_2^+$ )
- Nozzle: 1 mm  $\phi$ , Pierce

### Supersonic Expansion Discharge

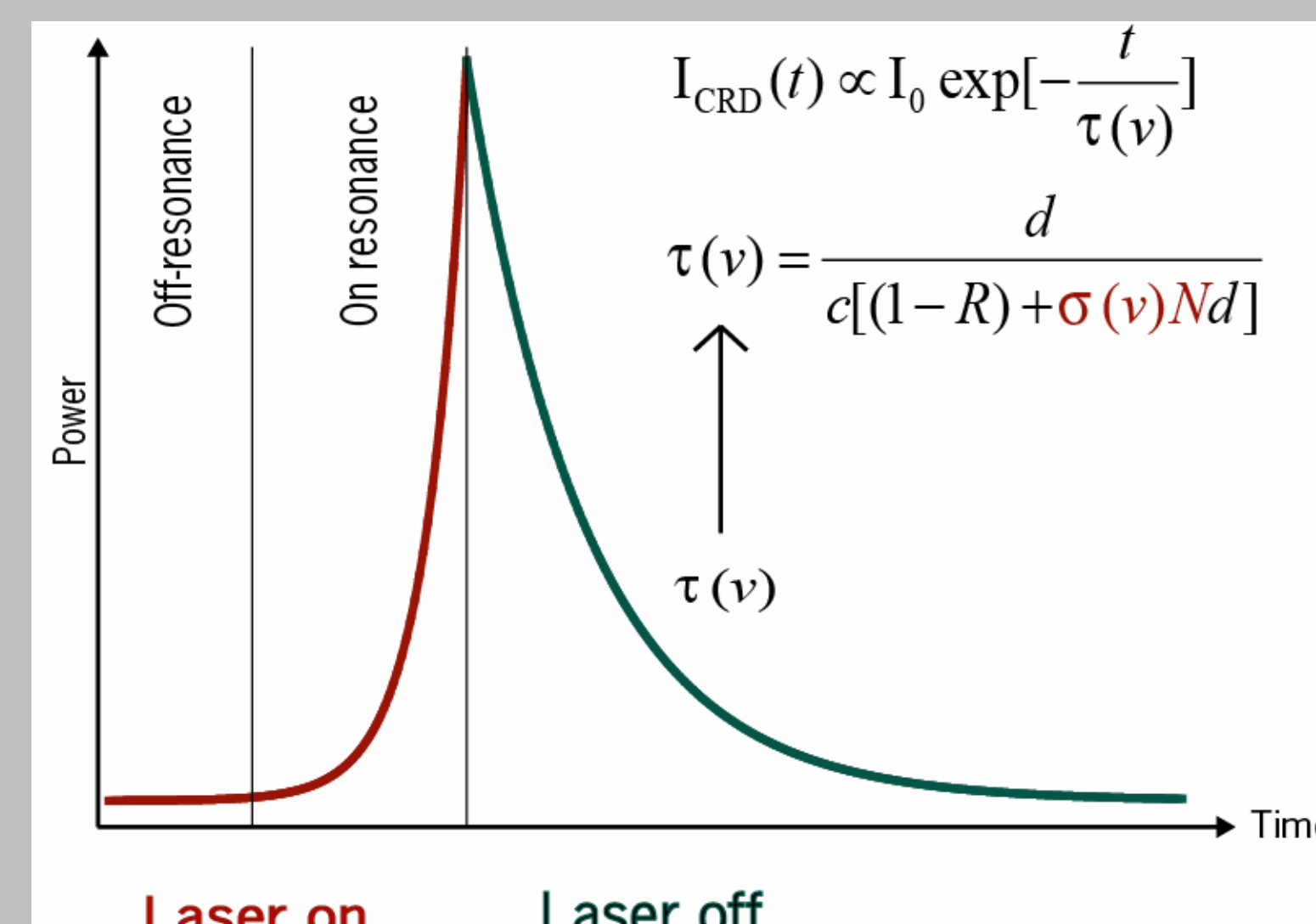
We are also developing and characterizing a continuous supersonic expansion discharge source for producing rotationally cold molecular ions.



## Spectroscopic Techniques

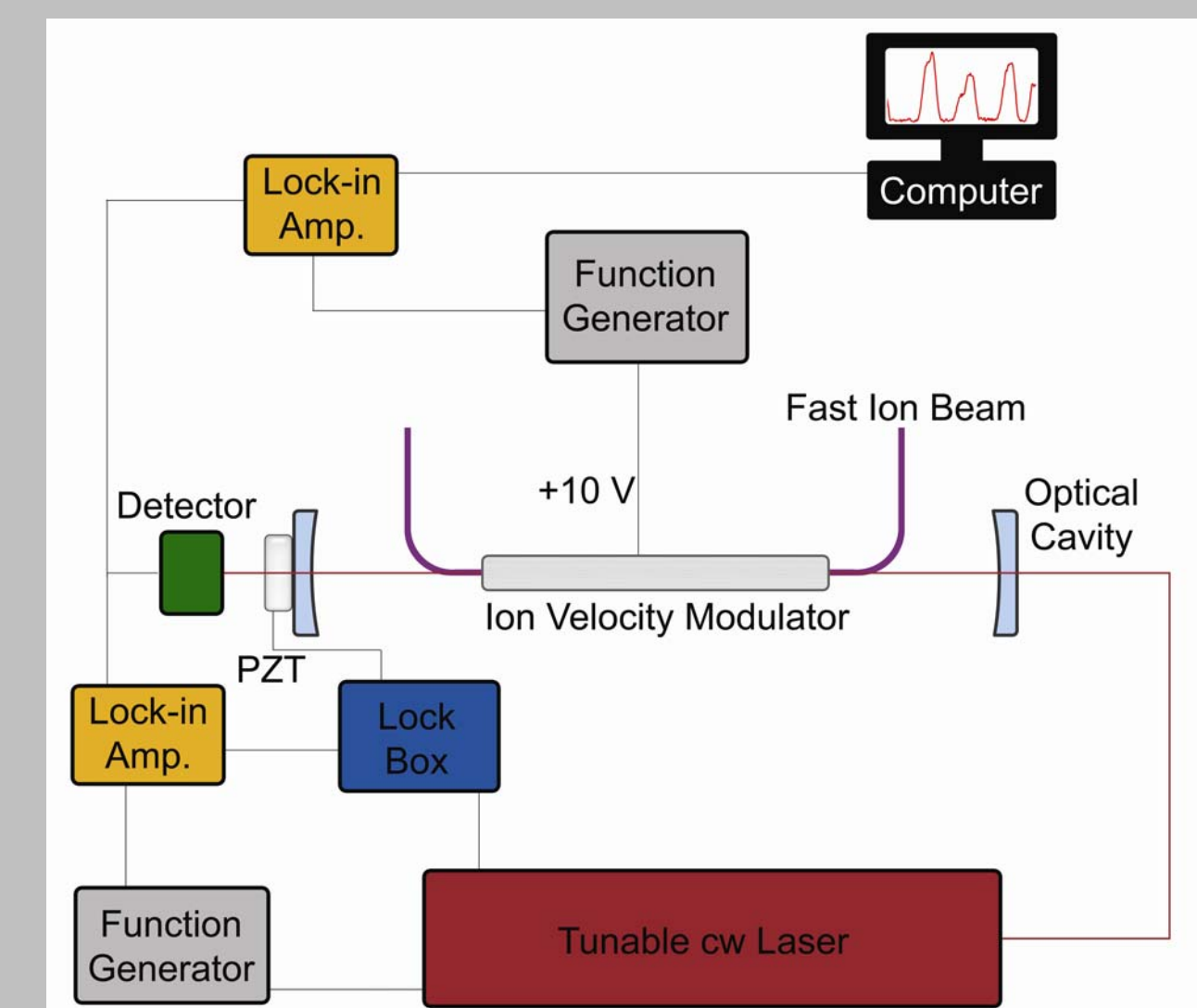
### Continuous-Wave Cavity Ringdown

- Easy to implement (AOM, comparator)
- Min. detectable absorption  $(\Delta I/I)_{\text{min}} \sim 10^{-7}$



### Cavity Enhanced Beam Velocity Modulation

- More involved (locking loop)
- $(\Delta I/I)_{\text{min}} \sim 10^{-7}$  / Finesse

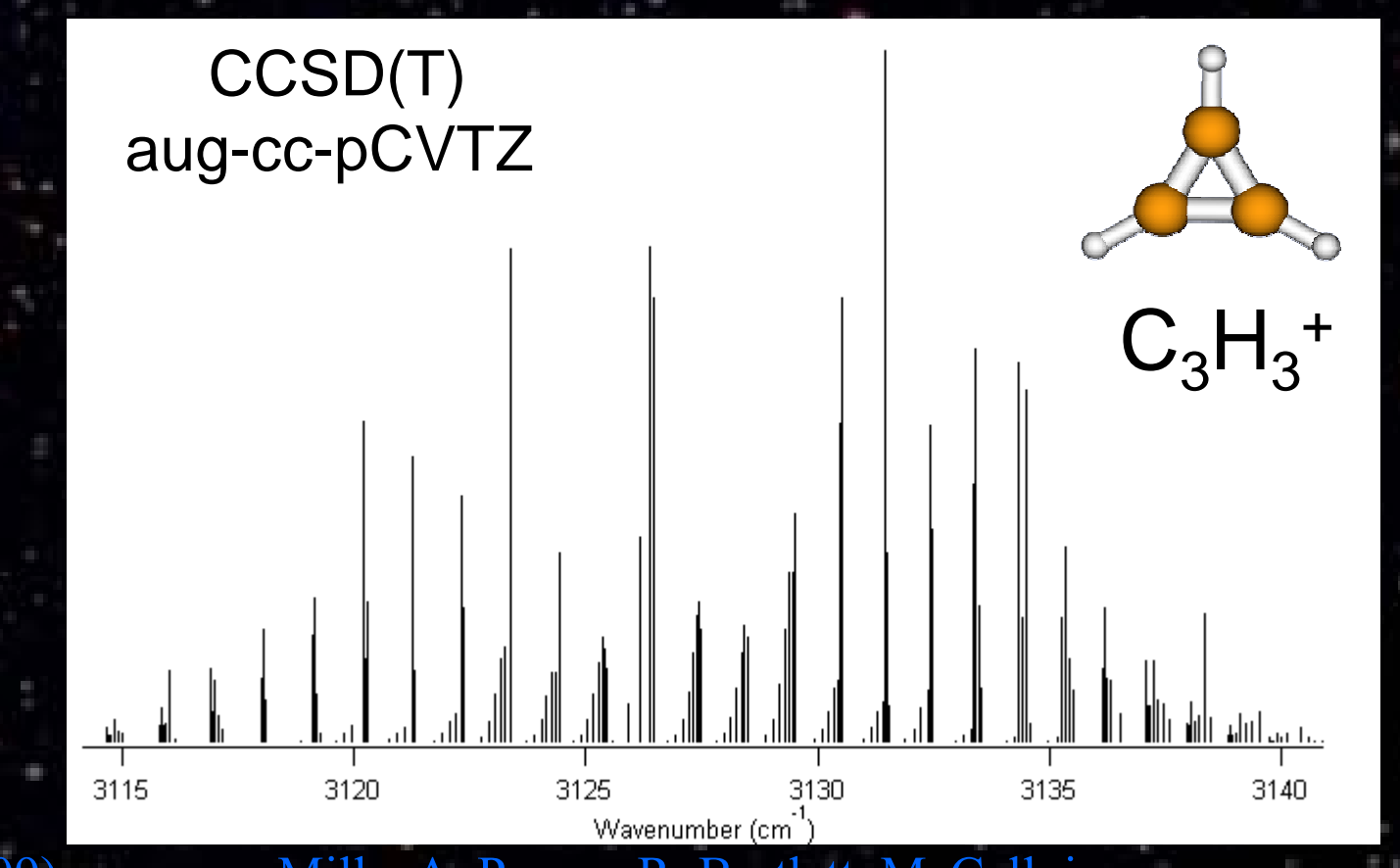
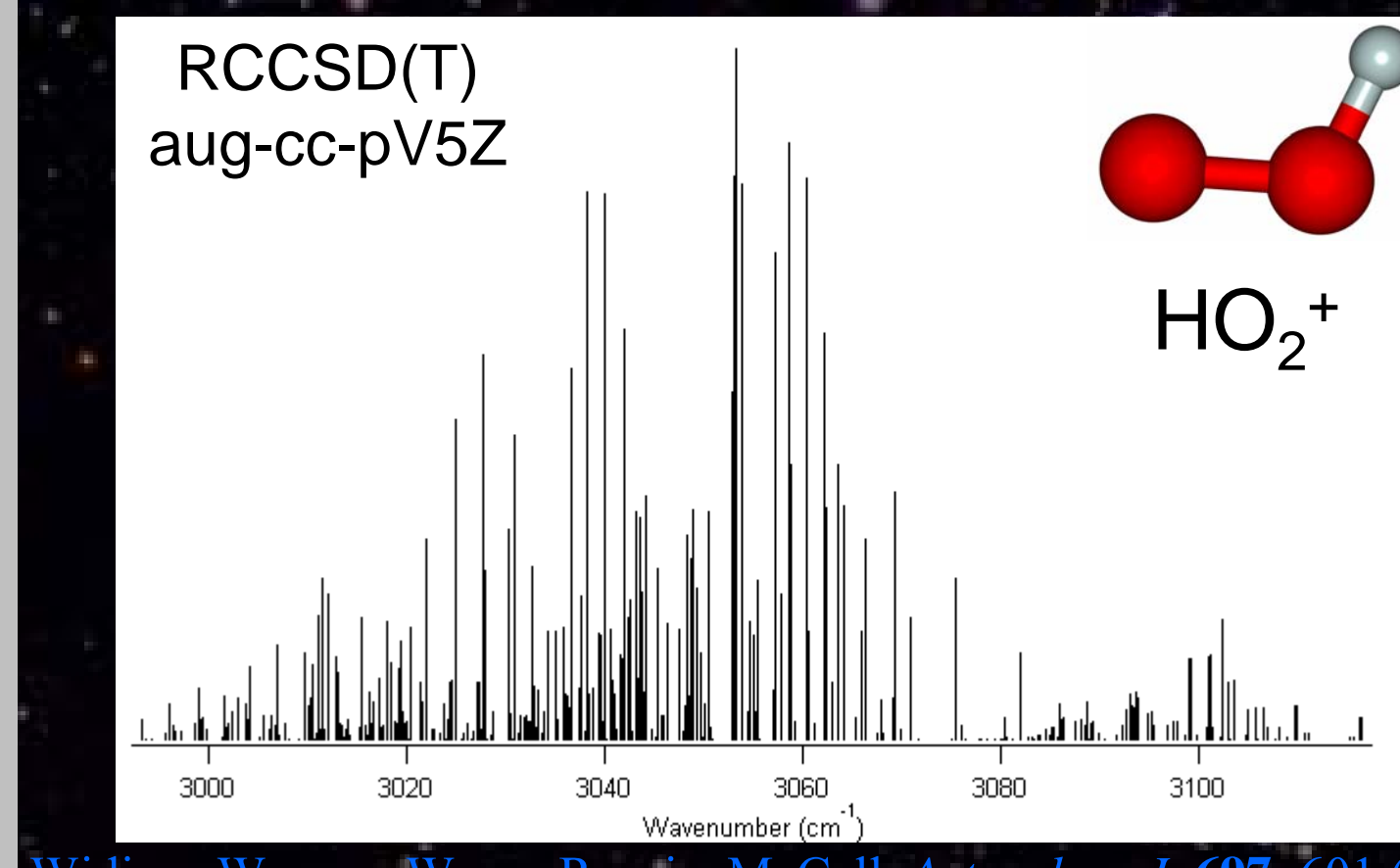


## Target Molecular Ions

We plan to study larger and more complex ions than can be studied in hot plasmas, including  $\text{N}_5^+$ ,  $\text{H}_5\text{O}_2^+$ ,  $\text{CH}_5^+$ ,  $\text{C}_2\text{H}_5^+$ ,  $\text{C}_3\text{H}_3^+$ , and  $\text{C}_6\text{H}_7^+$ . We welcome input on what will be the best targets for supporting the Air Force mission!

## Theoretical Support

We have been collaborating with theorists to perform high-level *ab initio* calculations to help guide our experiments.



Widicus Weaver, Woon, Ruscic, McCall, *Astrophys. J.* 697, 601 (2009)

Mills, A. Perera, R. Bartlett, McCall, in prep